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There are different senses of entropy*:

- Thermodynamic Sense
- Information Sense
- Statistical Sense
- Disorder Sense
- Homogeneity Sense

Especially the "disorder" and "homogeneity" senses are related to and even require the notion/specification/definition of space

Only few formulas for Entropy comprise spatial aspects/entities

One example for an entropy formula comprising spatial entities is the Bekenstein-Hawking entropy $\mathrm{S}_{\text {BH }}$ which in its dimensionless form* reads:
$\mathbf{S}_{\text {Bн }}$ : dimensionless
Bekenstein-Hawking entropy

1/4: important factor

A: area of the event horizon/ surface area of the black hole
$\mathrm{L}_{\mathrm{p}}$ : Planck length/ a small positive number having the dimension of a length

In spite of describing a physics object - a black hole - having mass, charge spin etc. this formula only contains geometric entities

Objective of the presentation is to derive the structure of this formula based on geometric considerations.

## The Heaviside function

The approach starts from the Heaviside function* $\Theta\left(x_{0}\right)$ :

which can be used to describe a sphere or any other geometric object.
The volume $V$ of a sphere with radius $r_{0}$ in spherical coordinates is then given by:

$$
V=\iiint \Theta\left(r-r_{0}\right) r^{2} d r d \Omega=\frac{4}{3} \pi r_{0}^{3}
$$

with $\mathrm{d} \Omega$ being the differential solid angle: $\sin \Theta d \Theta d \varphi=d \Omega$

## The $\delta$ function

The Dirac $\delta$ function actually is defined* as the distributional derivative of the Heaviside function $\Theta(x)$ as

$$
\delta(x):=\frac{d \Theta(x)}{d x}
$$

Using the $\delta$ function the surface A of the sphere - and also the surface of more complex geometric objects - can easily be calculated:

$$
A=\iiint \delta\left(r-r_{0}\right) r^{2} d r d \Omega=4 \pi r_{0}^{2}
$$

This approach thus has allowed to calculate the area "A" as the first step towards deriving the entropy of a geometric sphere

$$
S_{G S}=\frac{A}{4 l_{p}^{2}}
$$

In fact, however, nothing has been said by now about entropy.
The next steps will have a closer look at the transition region of the Heaviside function and introduce the phase-field function

## The phase field description of a transition



The Heaviside function varies discontinuously from 1 to 0 in an infinitesimally small transition region. Nothing is thus known about the shape of this function in the transition region.

The phase- field variable $\Phi$ in contrast varies continuously from 1 to 0 in the transition region with finite width $\eta$

The shape of the transition in phase-field models depends on the choice of the potential. A double-well potential e.g. leads to a hyperbolic- tangent profile while a double obstacle potential leads to a cosine profile of the $\Phi$ function

However, nothing is a priori known about the shape of this function in the transition region also in phase-field models.

## Can we learn more about the interface region?



The Phase-field function $\Phi$ can be considered as a contineous formulation of the Heaviside function $\Theta$ if the interface thickness $\eta$ becomes infinitesimally small.

Is there a rationale for the shape of both the Heaviside and the phase-field functions in the transition region?

## Entropy of a single interface layer: the Jackson Model



The Jackson model*:

- is used to describe facetted growth of crystals
- assumes ideal mixing of the two states (solid/liquid) in a single interface layer between the bulk states
- describes the entropy of the interface as:

$$
S=\Phi \ln \Phi+(1-\Phi) \ln (1-\Phi)
$$

- which generates $\Phi=0.5$ as the most probable value


The Kossel model*:

- is a discrete model
- is used to describe the growth of crystals with diffuse interfaces
- assumes attachment of solid on existing solid only (no overhang)
- describes a stepwise transition from $100 \%$ solid (the 4 left layers) to $100 \%$ liquid (from layer 11 to the right)
- provides the basis for Temkin's discrete formulation of the entropy of a diffuse interface


## Entropy of a diffuse interface: the Temkin Model



## The Temkin model*:

- is used to describe growth of crystals with diffuse interfaces
- assumes ideal mixing between two adjacent states/layers in a multilayer interface
- describes the entropy of the diffuse interface as:

$$
S=-\sum_{n=-\infty}^{\infty}\left(\Phi_{n-1}-\Phi_{n}\right) \ln \left(\Phi_{n-1}-\Phi_{n}\right)
$$

- recovers the Jackson model as a limiting case for a single interface layer


## Highlighting the importance of the Temkin model

## The Temkin model:



The gradient in the
Temkin model is identified as follows:

- introduces neighborhood relations between adjacent layers and thus an "order" resp. „disorder" sense
- introduces a gradient and thus a length scale into the formulation of entropy
- can be extended to a continuous formulation
- can be extended to 3 dimensions

$$
d \Phi_{n}=\Phi_{n-1}-\Phi_{n}=\int_{n l}^{(n-1) l} \frac{d \Phi}{d r} d r=\frac{d \Phi_{n}}{d r} \int_{n l}^{(n-1) l} d r=l \frac{d \Phi_{n}}{d r}
$$

with „"" being the distance between two adjacent layers and the gradient being assumed as constant between these two layers

## From discrete to continuous*

$$
\begin{gathered}
r(n)=r_{0}+n l \text { and } d n=\frac{d r}{l} \\
S=-\sum_{n=-\infty}^{\infty}\left(\Phi_{n-1}-\Phi_{n}\right) \ln \left(\Phi_{n-1}-\Phi_{n}\right)=-\sum_{n=-\infty}^{\infty}\left\{l \frac{d \Phi}{d r}(n l)\right\} \ln \left\{l \frac{d \Phi}{d r}(n l)\right\}
\end{gathered}
$$

Making the transition from discrete to continuous:

$$
-\sum_{n=-\infty}^{\infty}\left\{l \frac{d \Phi}{d r}(n l)\right\} l n\left\{l \frac{d \Phi}{d r}(n l)\right\} \rightarrow-\int_{-\infty}^{\infty}\left\{l \frac{d \Phi}{d r}(n l)\right\} \ln \left\{l \frac{d \Phi_{n}}{d r}(n l)\right\} d n
$$

and substituting: $\quad n l=r-r_{0}$ and $d n=\frac{d r}{l}$
in 1 dimension yields: $\quad S=-\int_{-\infty}^{\infty}\left\{l \nabla_{r} \Phi\left(r-r_{0}\right)\right\} \ln \left\{l \nabla_{r} \Phi\left(r-r_{0}\right)\right\} \frac{d r}{l}$

## Extending to 3D

extending the formulation to 3 dimensions in cartesian coordinates reads:

$$
S=-\iiint_{-\infty}^{\infty}(\overrightarrow{\mathrm{l}} \vec{\nabla} \phi) \ln (\overrightarrow{\mathrm{l}} \vec{\nabla} \phi) \frac{d x}{l_{x}} \frac{d y}{l_{y}} \frac{d z}{l_{z}}
$$

Assuming isotropy of space resp. of the discretization i.e. $l_{x}=l_{y}=l_{z}=l_{p}$ eventually leads to

$$
S=-\iint_{-\infty}^{\infty} \int_{-\infty}^{(\vec{I} \vec{\nabla} \phi) \ln (\overrightarrow{I \nabla} \phi)} \frac{l_{p}^{3}}{l} d x d y d z
$$

The term

$$
s=\frac{(\overrightarrow{\mathrm{l}} \vec{\nabla} \phi) \ln (\overrightarrow{\mathrm{l}} \vec{\nabla} \phi)}{l_{p}^{3}}
$$

can be interpreted as an entropy density.

## Extending to 3D in spherical coordinates

$$
S=-\iiint_{-\infty}^{\infty}(\overrightarrow{\mathrm{l}} \vec{\nabla} \phi) \ln (\overrightarrow{\mathrm{l}} \vec{\nabla} \phi) \frac{d x}{l_{x}} \frac{d y}{l_{y}} \frac{d z}{l_{z}}
$$

Switching to spherical coordinates yields:

$$
\begin{gathered}
\frac{d x}{l_{p}} \frac{d y}{l_{p}} \frac{d z}{l_{p}}=\frac{1}{l_{p}^{3}} r^{2} d r \sin \Theta d \Theta d \varphi=\frac{r^{2}}{l_{p}^{2}} \frac{d r}{l_{p}} d \Omega \\
S=-\iiint(\overrightarrow{\mathrm{l}} \phi(\mathrm{r})) \ln (\overrightarrow{\mathrm{l}} \vec{\nabla} \phi(\mathrm{r})) r^{2} \frac{d r}{l_{p}} \frac{d \Omega}{l_{p}^{2}}
\end{gathered}
$$

Assuming isotropy (i.e. no dependence on angular coordinates) allows for integration over the solid angle $\mathrm{d} \Omega$ :

$$
S=-\frac{4 \pi}{l_{p}^{2}} \int_{0}^{\infty}\left(\overrightarrow{\mathrm{l}} \vec{\nabla} \phi\left(\mathrm{r}-\mathrm{r}_{0}\right)\right) \ln \left(\overrightarrow{\mathrm{l}} \vec{\nabla} \phi\left(\mathrm{r}-\mathrm{r}_{0}\right)\right) r^{2} \frac{d r}{l_{p}}
$$

## Intermediate summary: the „Ip2" term

The integral

$$
S=-\frac{4 \pi}{l_{p}^{2}} \int_{0}^{\infty}\left(\overrightarrow{\mathrm{l}} \vec{\nabla} \phi\left(\mathrm{r}-\mathrm{r}_{0}\right)\right) \ln \left(\overrightarrow{\mathrm{l}} \vec{\nabla} \phi\left(\mathrm{r}-\mathrm{r}_{0}\right)\right) r^{2} \frac{d r}{l_{p}}
$$

will only deliver contributions at the interface $r=r_{0}$ as only at interfaces there is a finite gradient. The integrand can thus be considered being proportional to the $\delta$ function:

$$
\begin{gathered}
\frac{1}{l_{p}}\left(\overrightarrow{\mathrm{l}} \vec{\nabla} \Phi\left(\mathrm{r}-\mathrm{r}_{0}\right)\right) \ln \left(\overrightarrow{\mathrm{l}} \vec{\nabla} \Phi\left(\mathrm{r}-\mathrm{r}_{0}\right)\right)=\text { constant } * \delta\left(\mathrm{r}-\mathrm{r}_{0}\right) \\
S=-\frac{4 \pi}{l_{p}^{2}} \int_{0}^{\infty} \text { constant } * \delta\left(\mathrm{r}-\mathrm{r}_{0}\right) r^{2} d r=- \text { constant } * \frac{4 \pi r_{0}^{2}}{l_{p}^{2}}=- \text { constant } * \frac{A}{l_{p}^{2}}
\end{gathered}
$$

The entropy of a geometric sphere $\mathrm{S}_{\mathrm{GS}}$ thus gets closer to the formulation known for the Bekenstein-Hawking entropy $\mathrm{S}_{\text {BH }}$ of a black hole:

$$
S_{G S}=- \text { constant } * \frac{A}{l_{p}^{2}}=\frac{A}{4 l_{p}^{2}}
$$

Can we learn more about the shape of the transition?
Can we learn more from exploiting the term:
? $\quad \frac{1}{l_{p}}\left(\overrightarrow{I \nabla} \Phi\left(\mathrm{r}-\mathrm{r}_{0}\right)\right) \ln \left(\overrightarrow{\mathrm{I} \nabla \Phi}\left(\mathrm{r}-\mathrm{r}_{0}\right)\right)$
Is there a way to explain the factor $1 / 4$ ?
Statistics of „contrast" might help with „contrast" being defined as....

$$
\text { contrast }:=\vec{I} \vec{\nabla} \Phi
$$

## From average gradients to distribution of gradients (resp. contrast)



Possible shapes of the $\Phi$ function in the transition region:

A constant average gradient (blue) leads to an extremely narrow distribution of contrast being centered around $I_{p} / \eta$

The green shapes lead to high counts for small contrast

The red shape leads to a broader distribution of small and high contrast values

An entropy type distribution of contrast $x_{i}(i=10)$ :

$$
H(x)=-10^{*} x^{*} \ln (x)
$$

is indicated as the red-line overlay

## Averaging the distribution of contrast

The average of the contrast distribution can be calculated as follows

$$
\left\langle l_{p} \nabla \Phi\right\rangle=\frac{\int_{l_{p} \nabla \Phi_{\min }}^{l_{p} \nabla \Phi_{\max }}\left(l_{p} \nabla \Phi\right) \ln \left(\mathrm{l}_{p} \nabla \Phi\right) d\left(l_{p} \nabla \Phi\right)}{\int_{l_{p} \nabla \Phi_{\min }}^{l_{p} \nabla \Phi_{\max }} d\left(l_{p} \nabla \Phi\right)}
$$

The minimum contrast in the distribution has the value 0 while the maximum contrast is 1 with the maximum gradient then being $1 / /_{\mathrm{p}}$. This allows to fix the boundaries of the integral to 0 resp. 1. For these boundaries the integral in the denominator yields a value of 1 . The remaining integral

$$
\left\langle l_{p} \nabla \Phi\right\rangle=\int_{0}^{1}\left(l_{p} \nabla \Phi\right) \ln \left(l_{p} \nabla \Phi\right) d\left(l_{p} \nabla \Phi\right)
$$

according to a standard formula* interestingly yields

$$
\int_{0}^{1} x \ln (x) d x=1\left[\frac{\ln 1}{2}-\frac{1}{4}\right]-0\left[\frac{\ln 0}{2}-\frac{1}{4}\right]=-\frac{1}{4}
$$

Replacing the contrast distribution in the integral

$$
S=-\frac{4 \pi}{l_{p}^{2}} \int_{0}^{\infty}\left(\overrightarrow{\mathrm{l}} \vec{\nabla} \phi\left(\mathrm{r}-\mathrm{r}_{0}\right)\right) \ln \left(\overrightarrow{\mathrm{l}} \vec{\nabla} \phi\left(\mathrm{r}-\mathrm{r}_{0}\right)\right) r^{2} \frac{d r}{l_{p}}
$$

by its average

$$
\left\langle l_{p} \nabla \Phi\right\rangle=-\frac{1}{4} \operatorname{resp} . \quad\langle\nabla \Phi\rangle=-\frac{1}{4 l_{p}}=-\frac{1}{4} \frac{1}{l_{p}}=-\frac{1}{4} \nabla \Phi_{\max }
$$

leads to

$$
S=\frac{4 \pi}{l_{p}^{2}} \int_{0}^{\infty} \frac{1}{4} r^{2}\left|\overrightarrow{\nabla_{\max }} \phi\left(\mathrm{r}-\mathrm{r}_{0}\right)\right| d r
$$

and thus eventually to

$$
S \sim \frac{4 \pi}{l_{p}^{2}} \int_{0}^{\infty} \frac{1}{4} r^{2} \delta\left(\mathrm{r}-\mathrm{r}_{0}\right) d r=\frac{4 \pi r_{0}^{2}}{4 l_{p}^{2}}=\frac{A}{4 l_{p}^{2}}
$$

## Summary

The structure of the Bekenstein- Hawking formula for the dimensionless entropy of a black hole has been derived for a geometric sphere

The derivation is based only on geometric considerations
Key ingredient to the approach is a statistical description of the transition region in a Heaviside resp. phase-field function.

Based on the Temkin entropy of a diffuse interface gradients are introduced in form of scalar products into the formulation of entropy for this purpose.

This introduces a length scale into entropy and provides a link between the world of entropy type models and the world of Laplacian type models (see following slides)

Most interesting physics and new insights - e.g. on entropic gravity - may emerge when applying and exploiting the „contrast- concept" in more depth (see final slide).
"Contrast" may also be considered as the contrast between two quantummechanical states

## Entropy type equations

$$
S=k_{\mathrm{B}} \ln W \quad \text { Boltzmann entropy }
$$

$$
S=-k_{\mathrm{B}} \sum p_{i} \ln p_{i} \quad \text { Gibbs-Boltzmann entropy }
$$

$$
\mathrm{H}=-p \cdot \log _{2} p-(1-p) \cdot \log _{2}(1-p)
$$

$$
H(X)=-\sum_{i=1}^{n} p\left(x_{i}\right) \log p\left(x_{i}\right) . \begin{aligned}
& \text { Shannon entropy (binary) } \\
& \text { Shannon entropy } \\
& \text { (general) }
\end{aligned}
$$

$$
S=-k_{\mathrm{B}} \operatorname{Tr}(\hat{\rho} \log (\hat{\rho})), \quad \text { von Neumann entropy }
$$

$$
H_{\alpha}(X)=\frac{1}{1-\alpha} \log \left(\sum_{i=1}^{n} p_{i}^{\alpha}\right) \text { Rényi entropy }
$$

Incomplete list of models for a statistical/entropic description
of entities in physics and in information theory

Most of these models have a logarithmic term as a common ingredient.

None of these expression comprises gradients and/or Laplacian operators

$$
S_{q}\left(p_{i}\right)=\frac{k}{q-1}\left(1-\sum_{i} p_{i}^{q}\right) \quad \text { Tsallis entropy }
$$

## Laplacian type equations

$$
\begin{aligned}
& -\Delta u=f \quad \text { Poisson Equation } \\
& \Delta \Phi(\mathbf{r})=-\frac{\rho(\mathbf{r})}{\varepsilon} \quad \text { Coulomb Equation } \\
& \Delta \Phi(\mathbf{r})=4 \pi \cdot G \cdot \rho(\mathbf{r}) \quad \text { Newton Equation } \\
& i \hbar \frac{\partial}{\partial t} \Psi(\mathbf{r}, t)=\left[\frac{-\hbar^{2}}{2 \mu} \nabla^{2}+V(\mathbf{r}, t)\right] \Psi(\mathbf{r}, t) \\
& \begin{array}{l}
\frac{\partial \phi(\mathbf{r}, t)}{\partial t}=D \nabla^{2} \phi(\mathbf{r}, t) \quad \text { Diffusion Equation } \\
\square=\frac{\partial^{2}}{c^{2} \partial t^{2}}-\Delta \text { Wave Equation (operator) } \\
\alpha \varepsilon^{2} \partial_{t} \phi=\varepsilon^{2} \nabla^{2} \phi-f^{\prime}(\phi)-\frac{e_{0}}{h_{0}} h^{\prime}(\phi) u+\tilde{\eta}(\mathbf{r}, t) \\
\frac{\partial c}{\partial t}=D \nabla^{2}\left(c^{3}-c-\gamma \nabla^{2} c\right) \quad \text { Cahnase-field Equation } \\
\hline
\end{array} \\
& \hline
\end{aligned}
$$

Incomplete list of models for a spatio-temporal description of stationary solutions or for the evolution in physics systems

Many of these models have a Laplacian operator as a common ingredient.

## Combining statistical and spatially resolved models



Bridging the gap between statistics/entropy type models and spatio-temporal models of the Laplacian world

## Benefits

First application of this concept:

## A Combined Entropy/Phase-Field Approach to Gravity

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․․ View Full-Text | Download PDF [1538 KB, uploaded 24 April 2017] | Browse Figures
resulted in :

- Poisson equation/Newtons law
- terms related to curvature of space,
- terms possibly explaining modified Newtonian dynamics

